

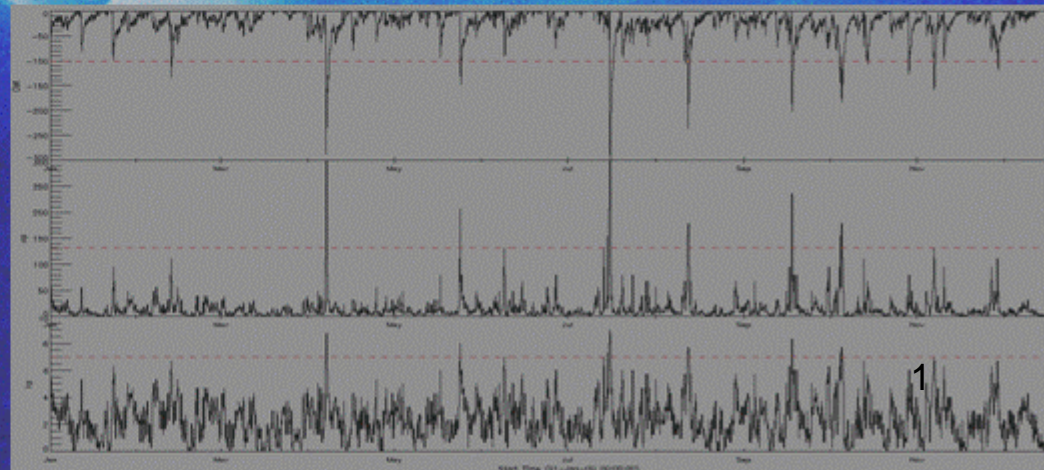
Oct. 16-21, 2005

Wintergreen, VA, USA

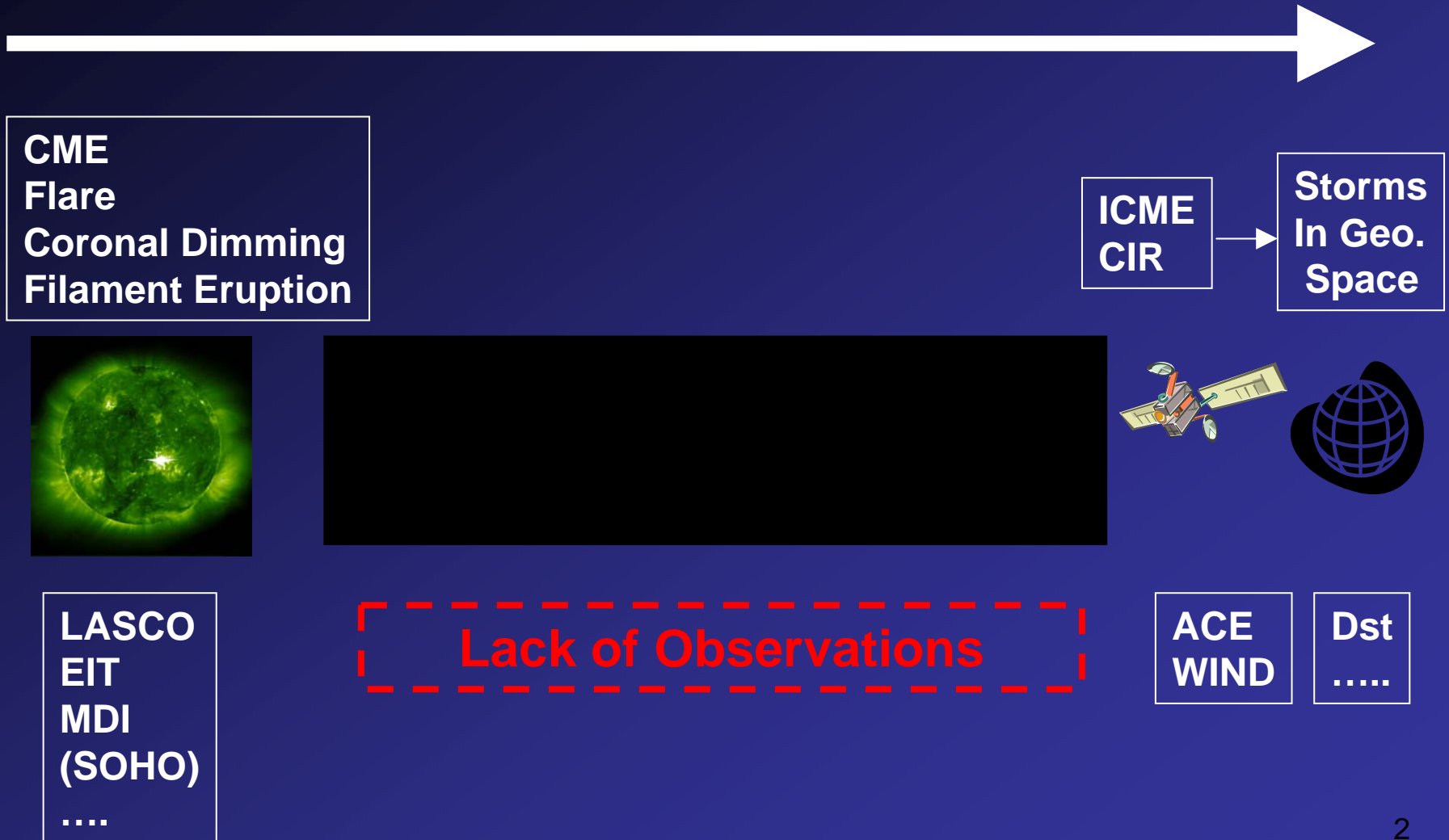
Solar and Space Physics and the vision for Space Exploration

The Sun-Earth Connection of Major Geomagnetic Storms

Jie Zhang
George Mason University

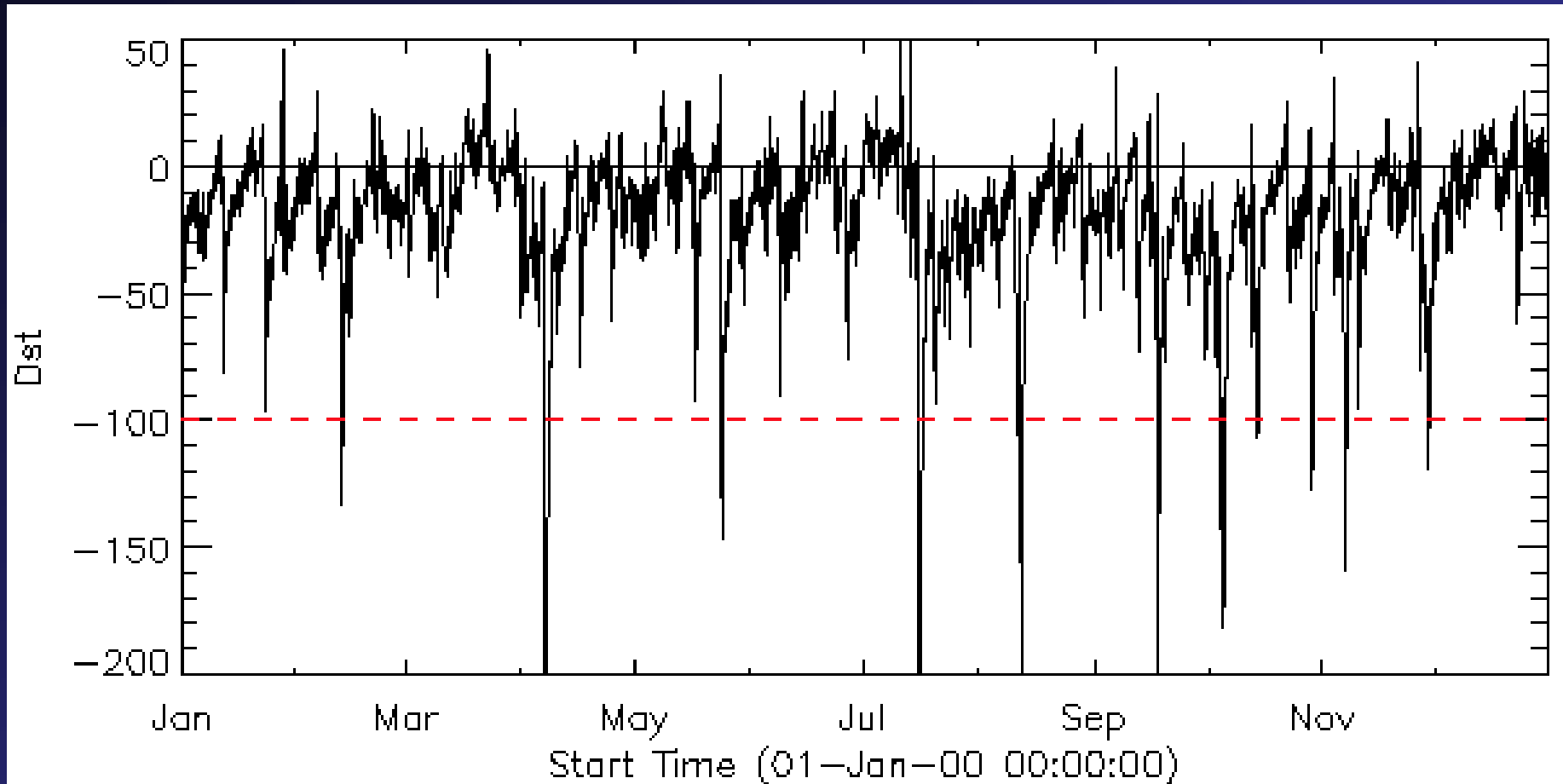


Event chain from Sun to Earth



Major Geomagnetic Storms

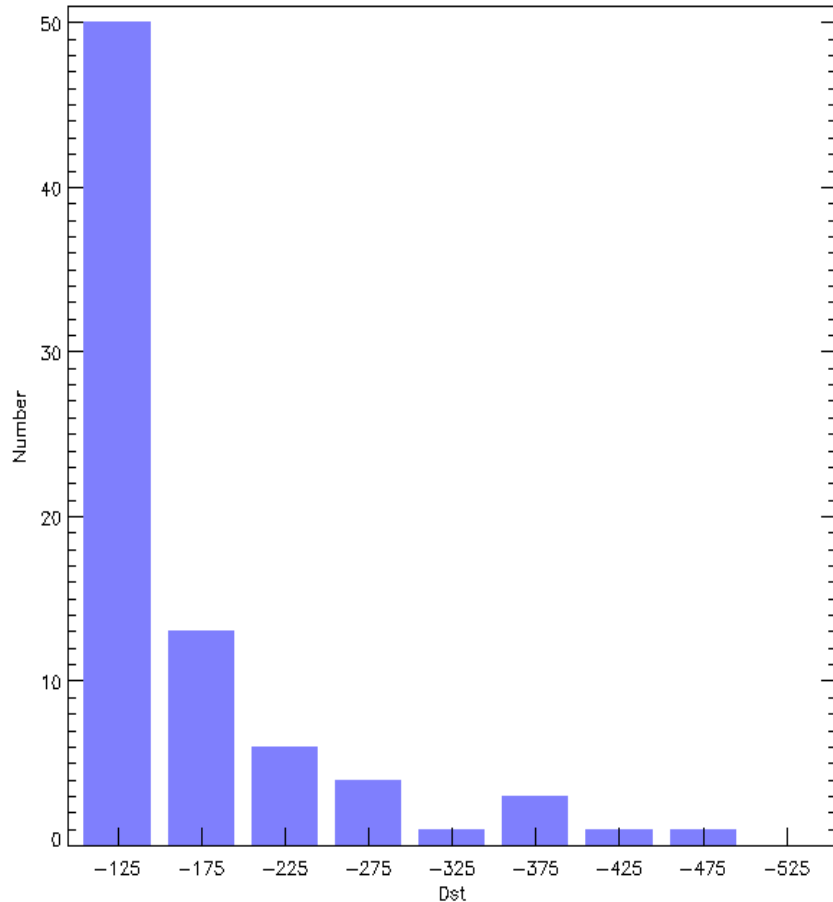
- $Dst \leq -100$



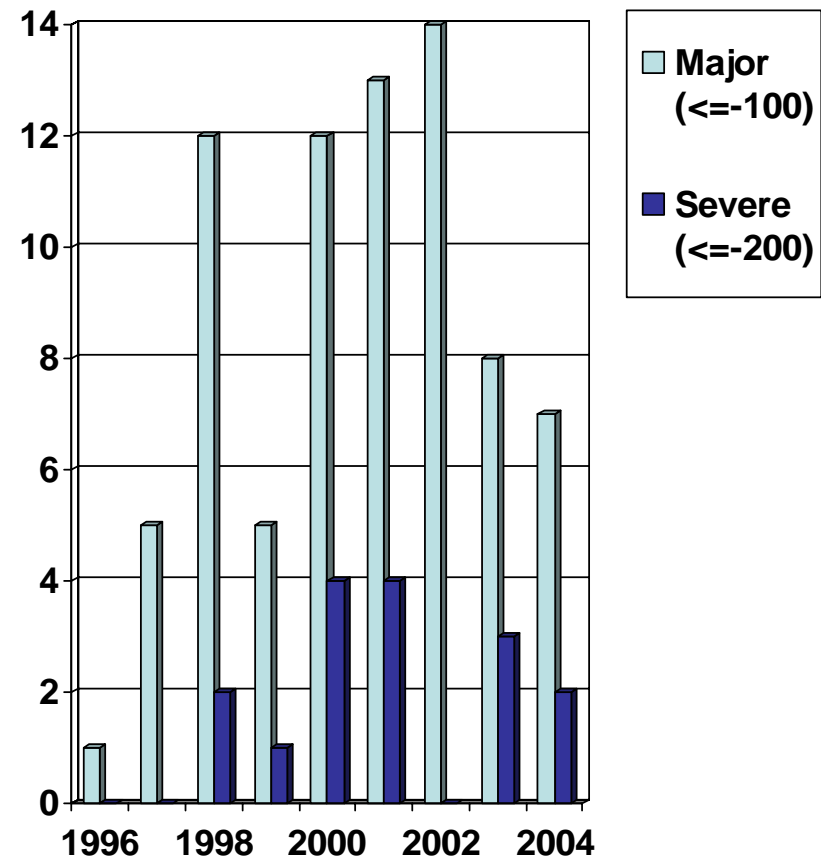
Dst index in 2000

Distribution of Major Storms (1996-2004)

- 79 major events ($Dst \leq -100$)
- 16 severe events ($Dst \leq -200$)



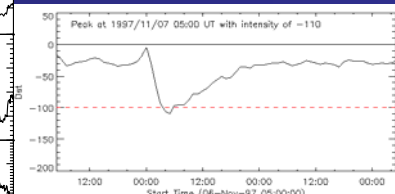
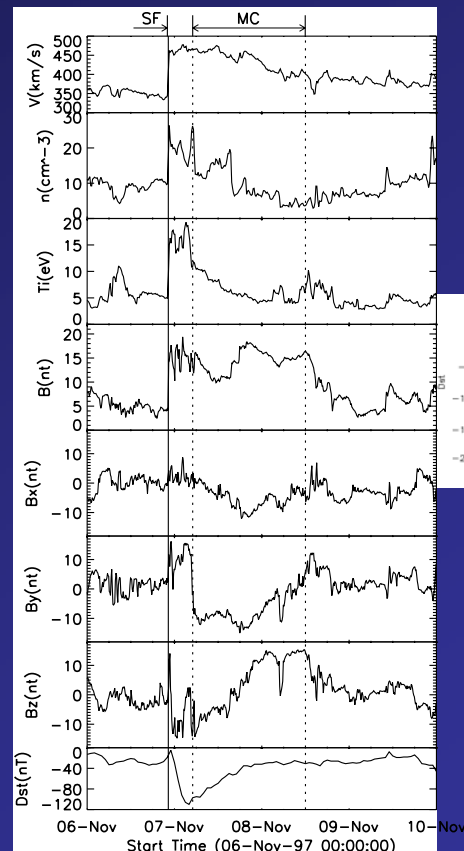
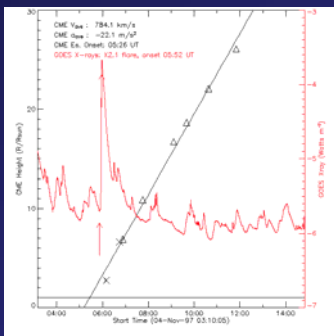
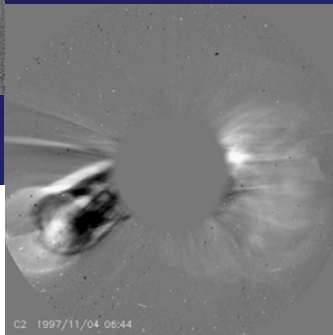
By Intensity



By Year

How to Find Solar Sources

- Step 1: set the 120-hr backward search window
- Step 2: find all halo CMEs in the window ($AW \geq 120$)
- Step 3: find out front-side halo CMEs, and locate their source region
- Step 4: find a reduced adaptive search window using solar wind velocity jump (Zhang et al. 2003)

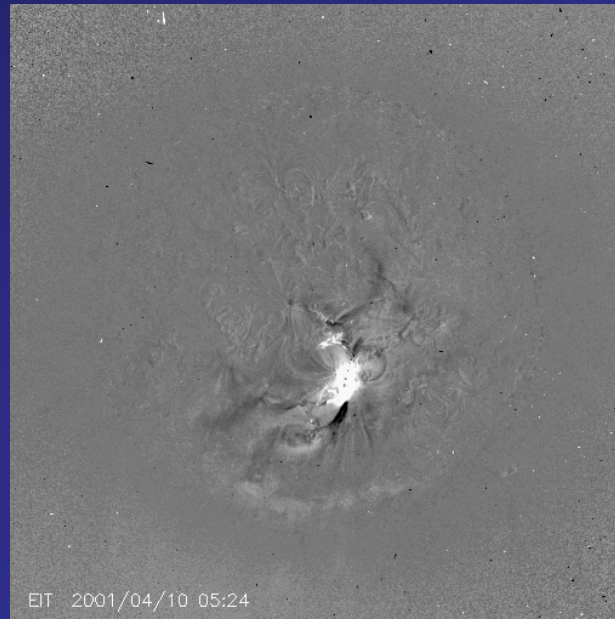
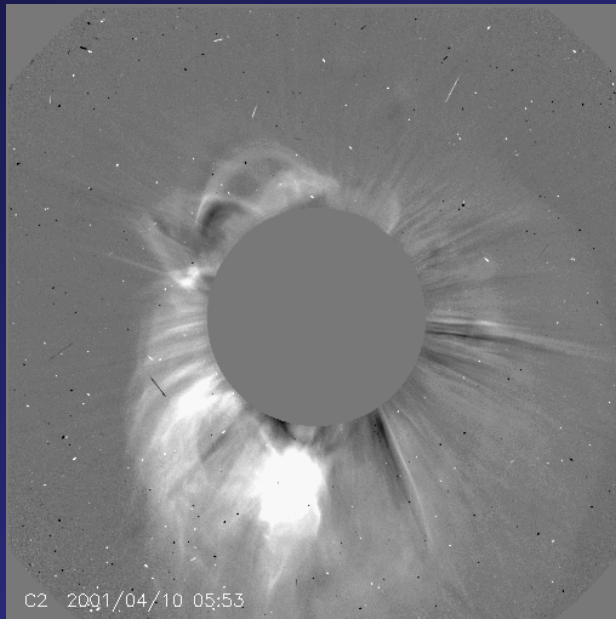
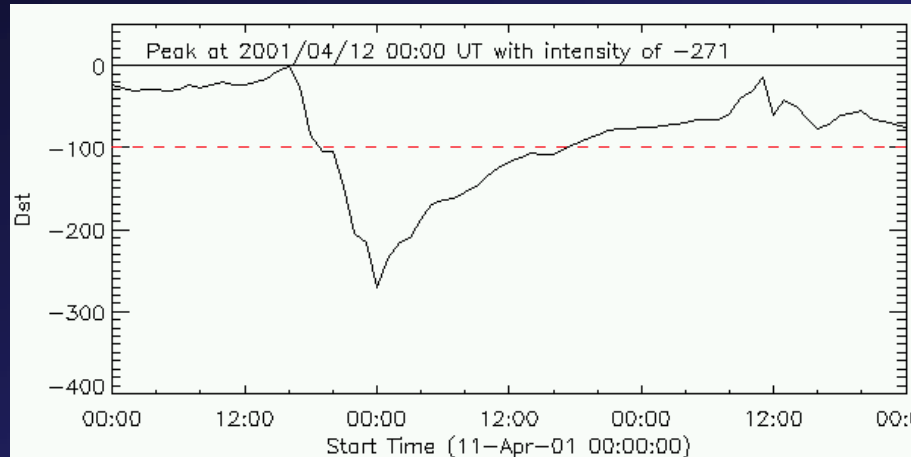


Three Types of Solar Drivers

1. **S Type: driven by a single CME**
2. **M Type: driven by multiple CMEs interacting in IP**
3. **C Type: driven by CIR from Coronal Hole**

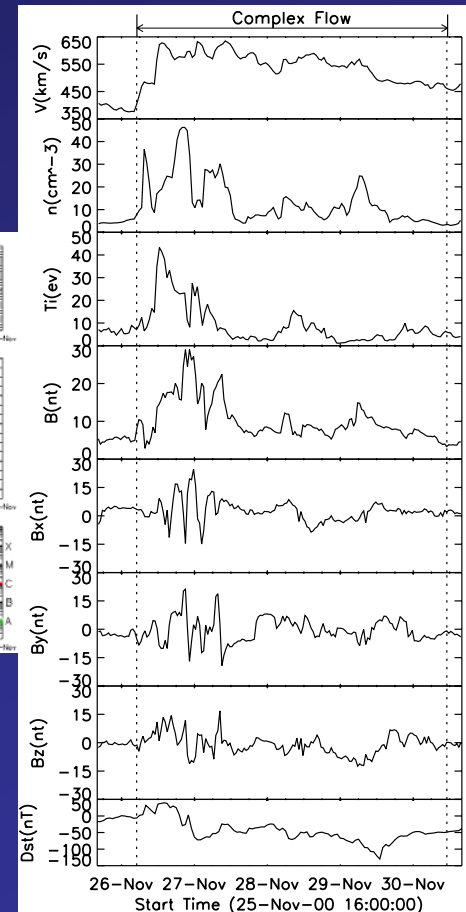
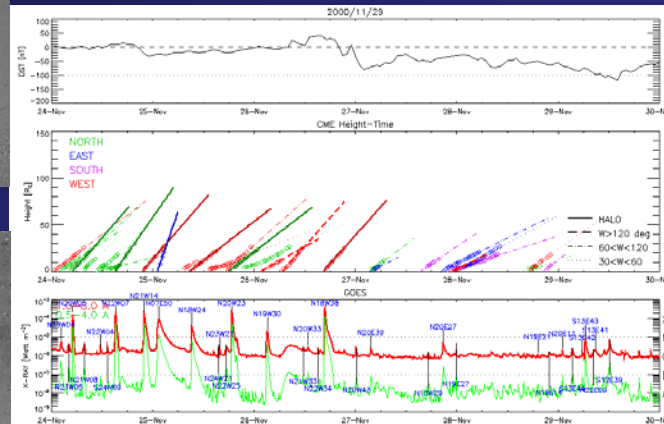
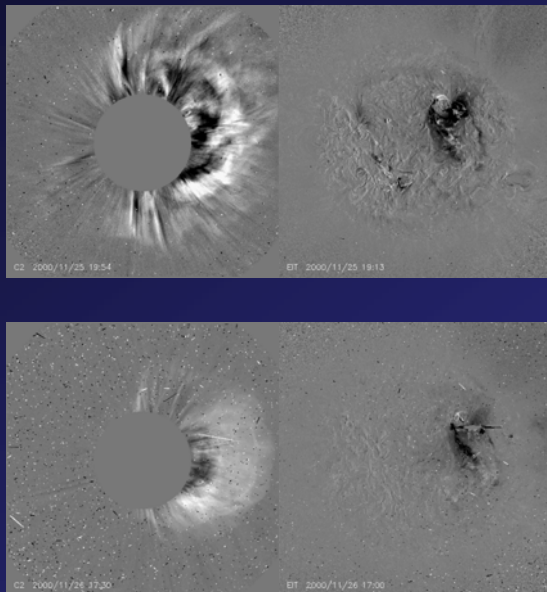
S Type Driver (example)

Dst peak at 2001/04/12 00:00 UT, driven by CME at 04/10 05:30 UT



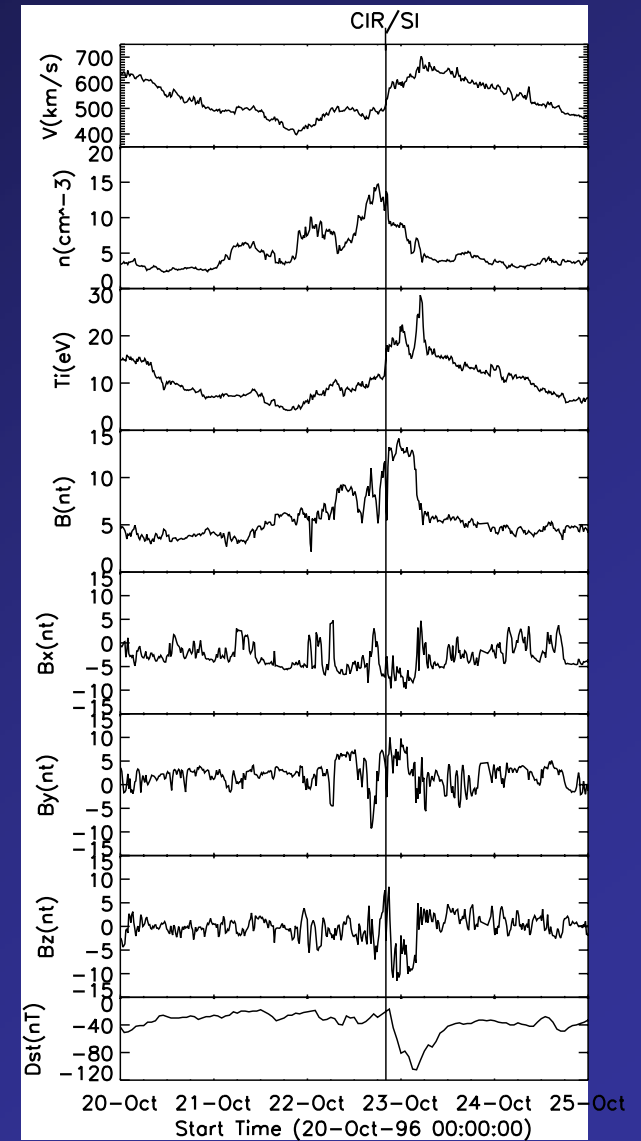
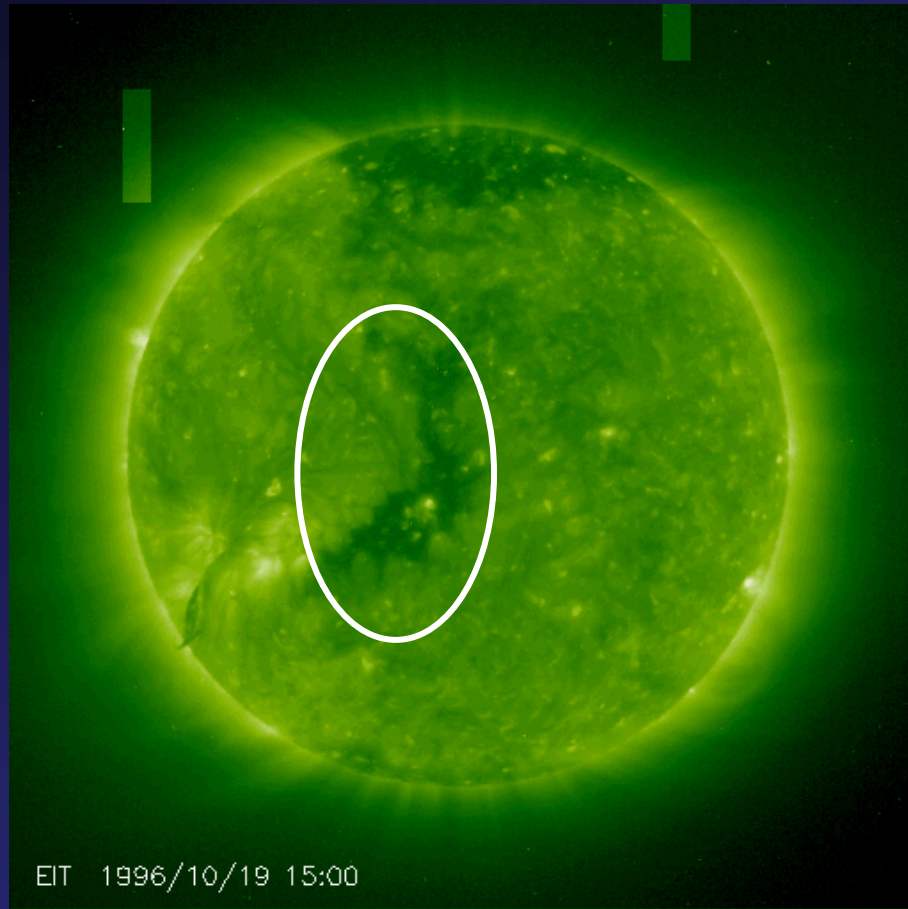
M Type Driver

- **Complex Dst plot, multiple CMEs (and flares)**
- **Complex solar wind flow**
- **Consecutive CMEs from same active region**

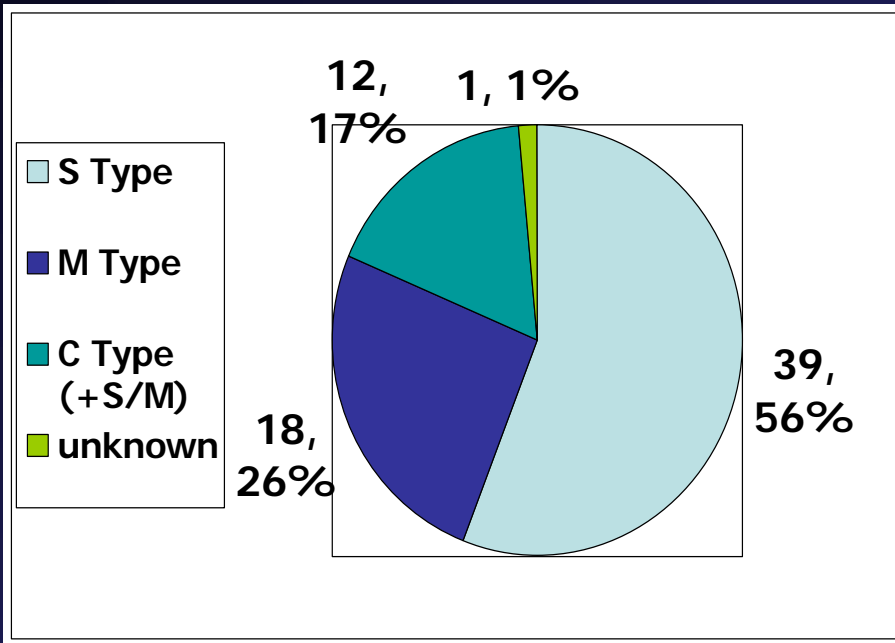


C Type Driver

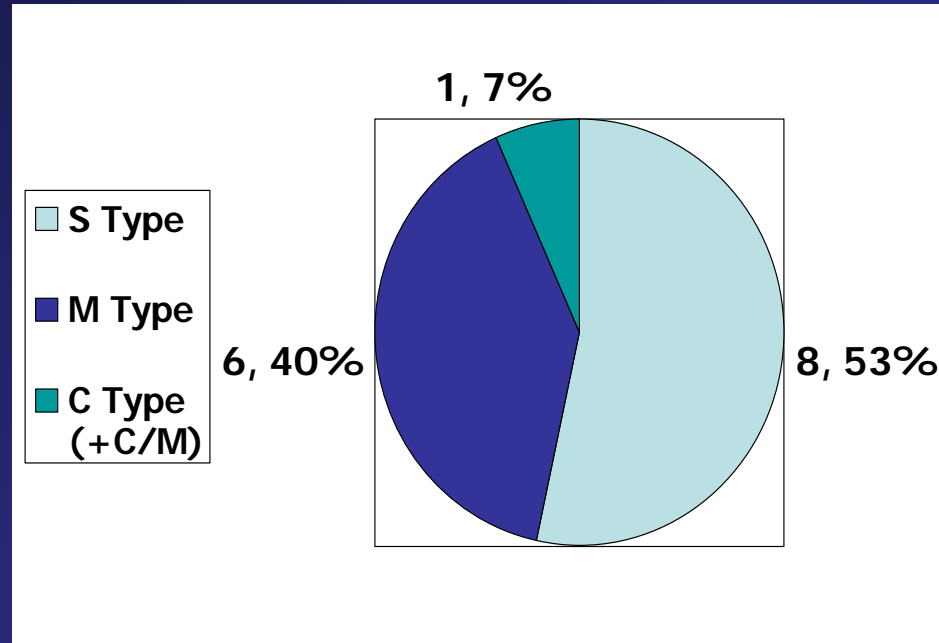
- CIR (Corotation Interaction Region)
- Coronal Hole



Solar Drivers of Major Storms

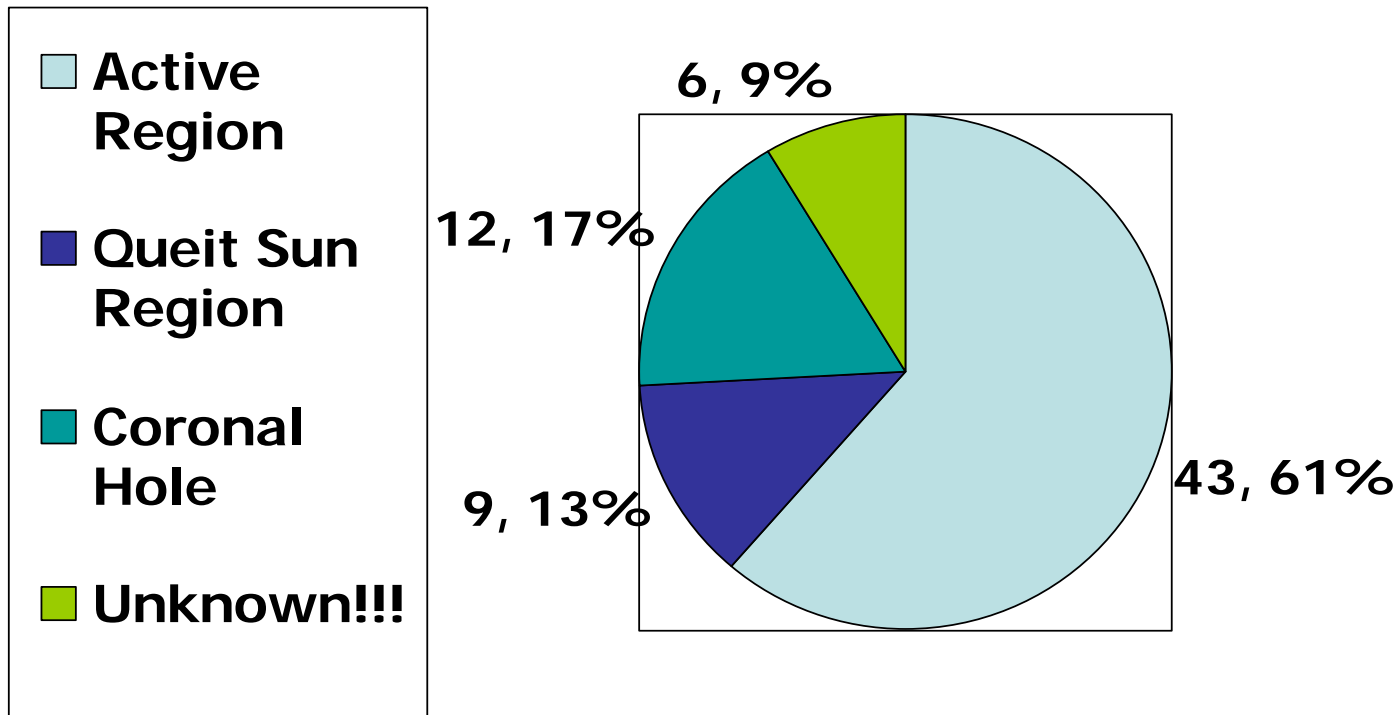


70 Major Storms
(excluding 9 in data gap)



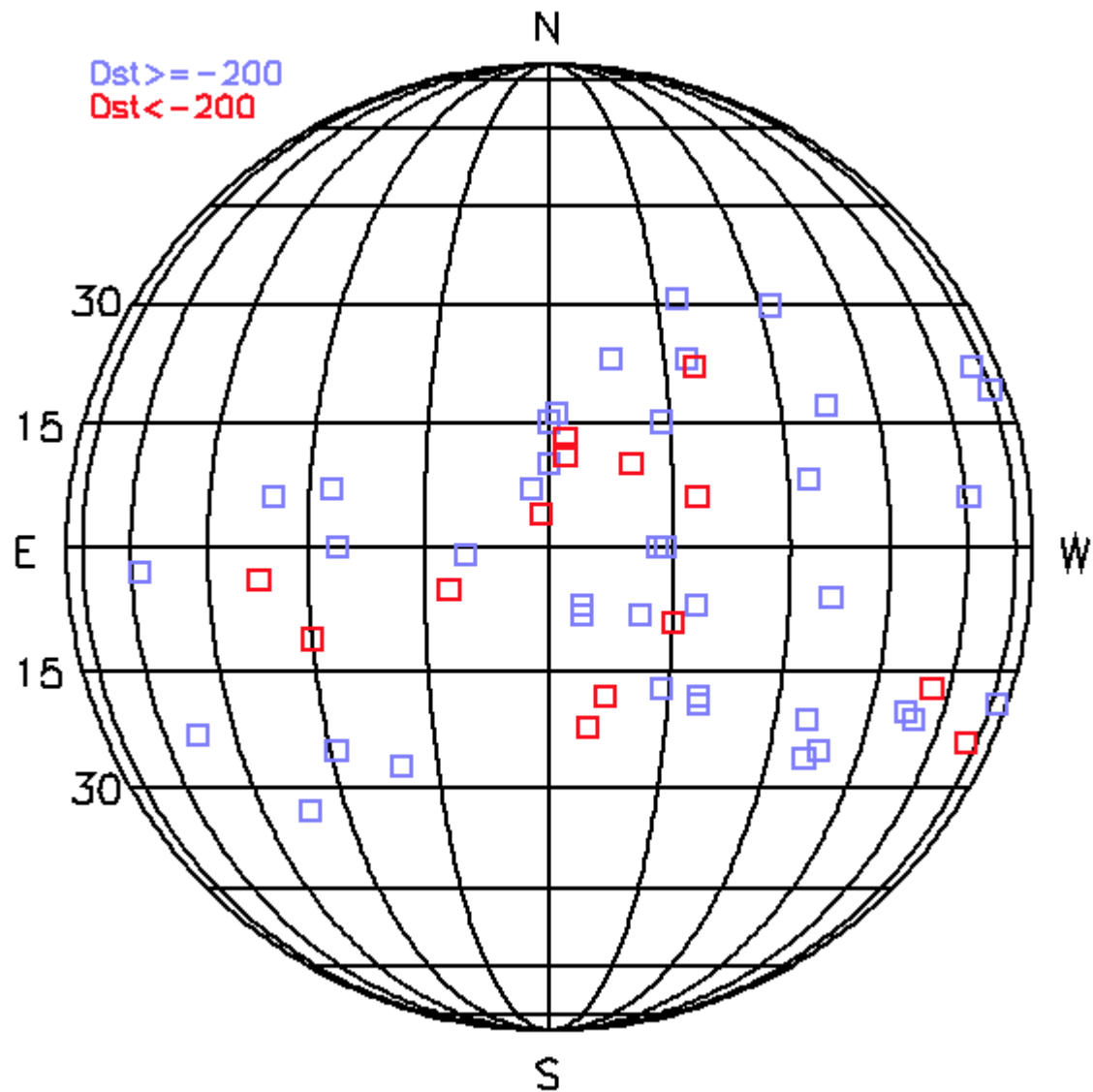
15 Severe Storms
(excluding 1 in data gap)

Source Regions of Major Storms



70 Major Storms

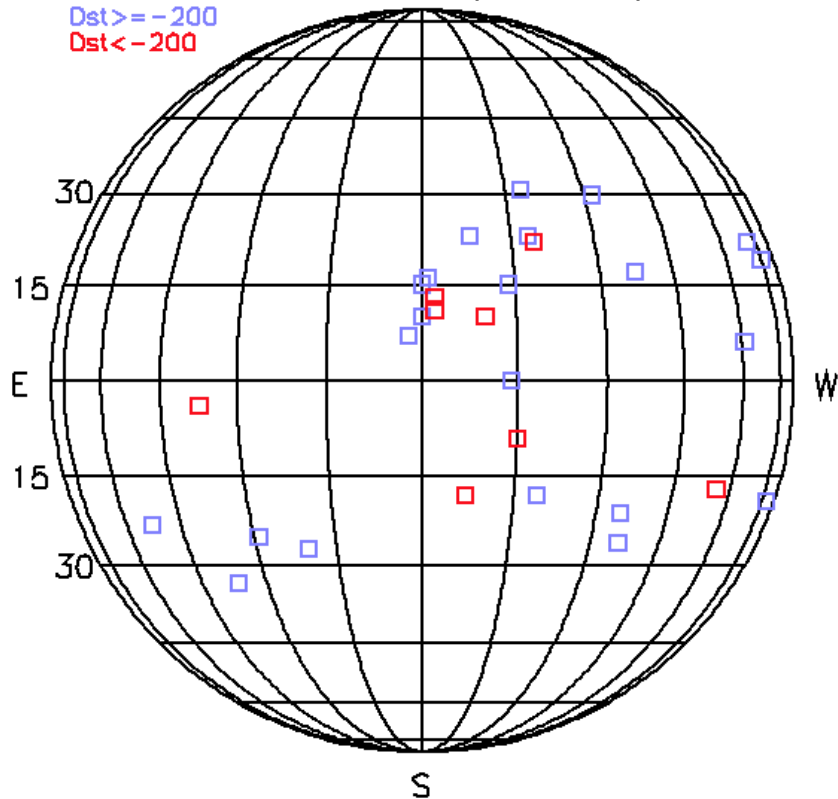
Source Region Distribution



Intra-Solar Cycle Variation of Source Regions

Before Reversal (Nov.2001)

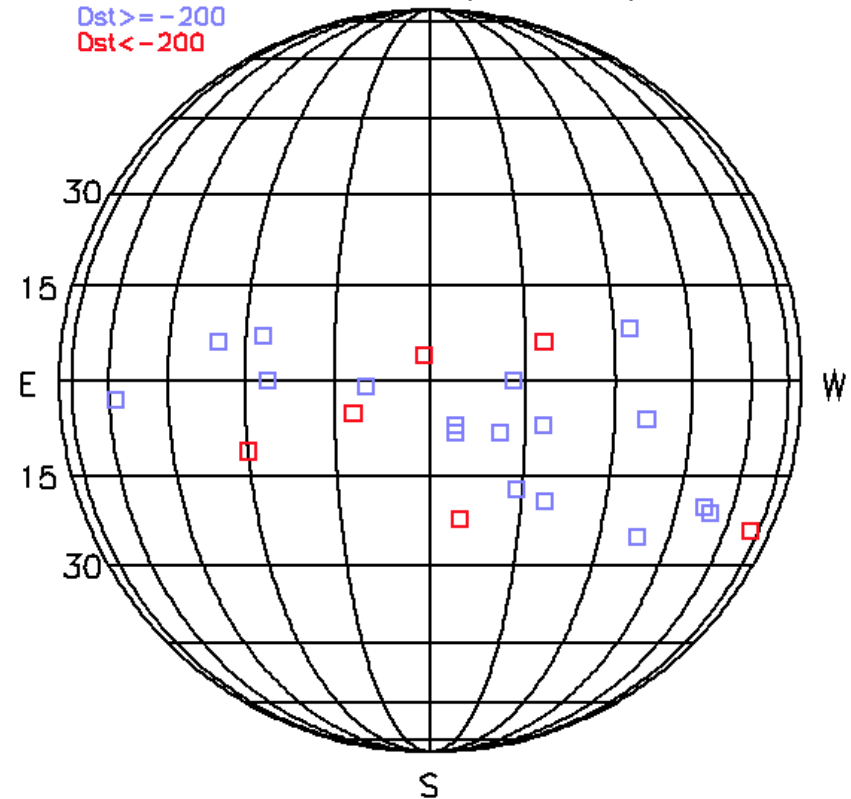
$Dst \geq -200$
 $Dst < -200$



Before Polar Reversal

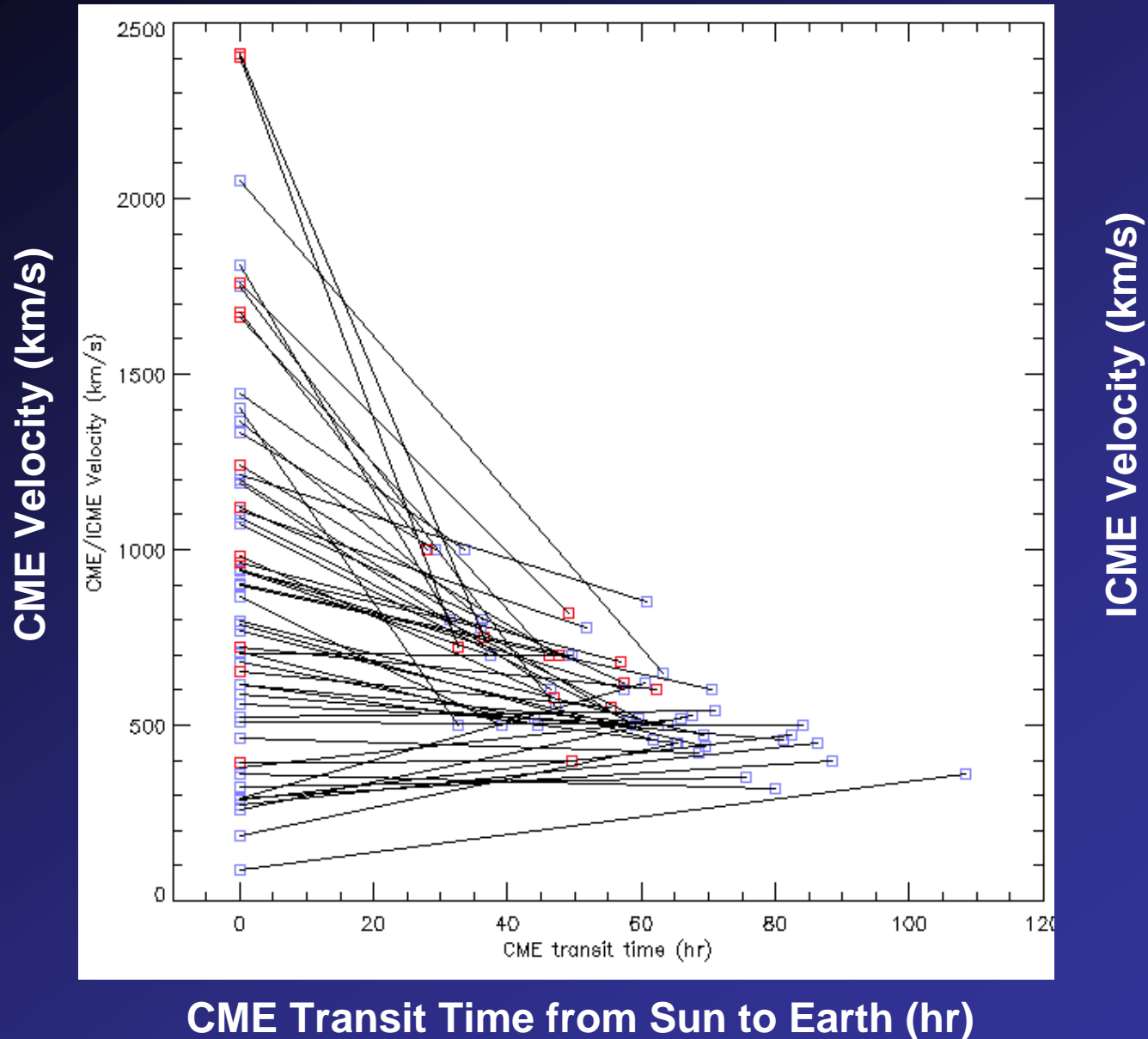
After Reversal (Nov.2001)

$Dst \geq -200$
 $Dst < -200$

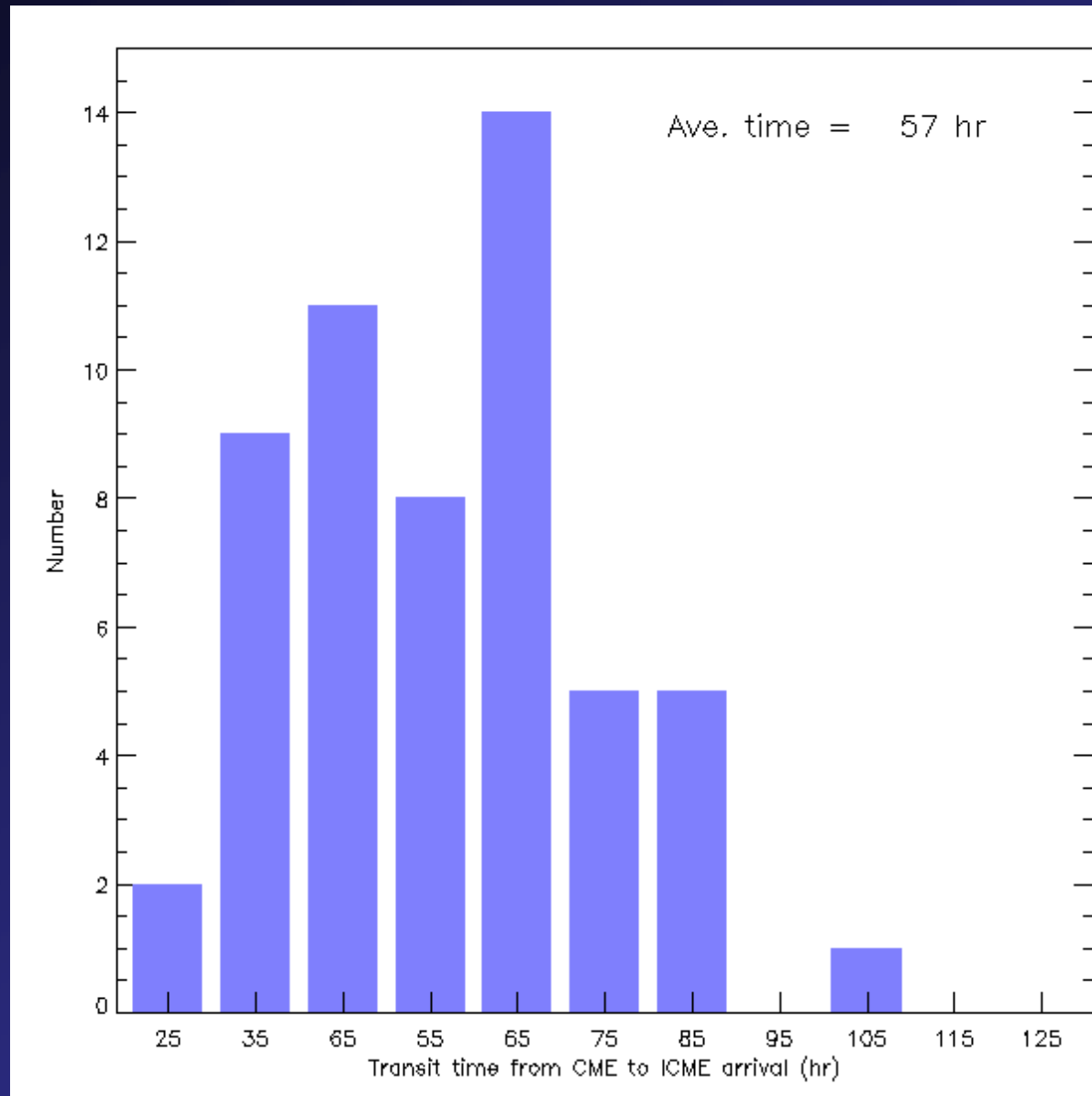


After Polar Reversal

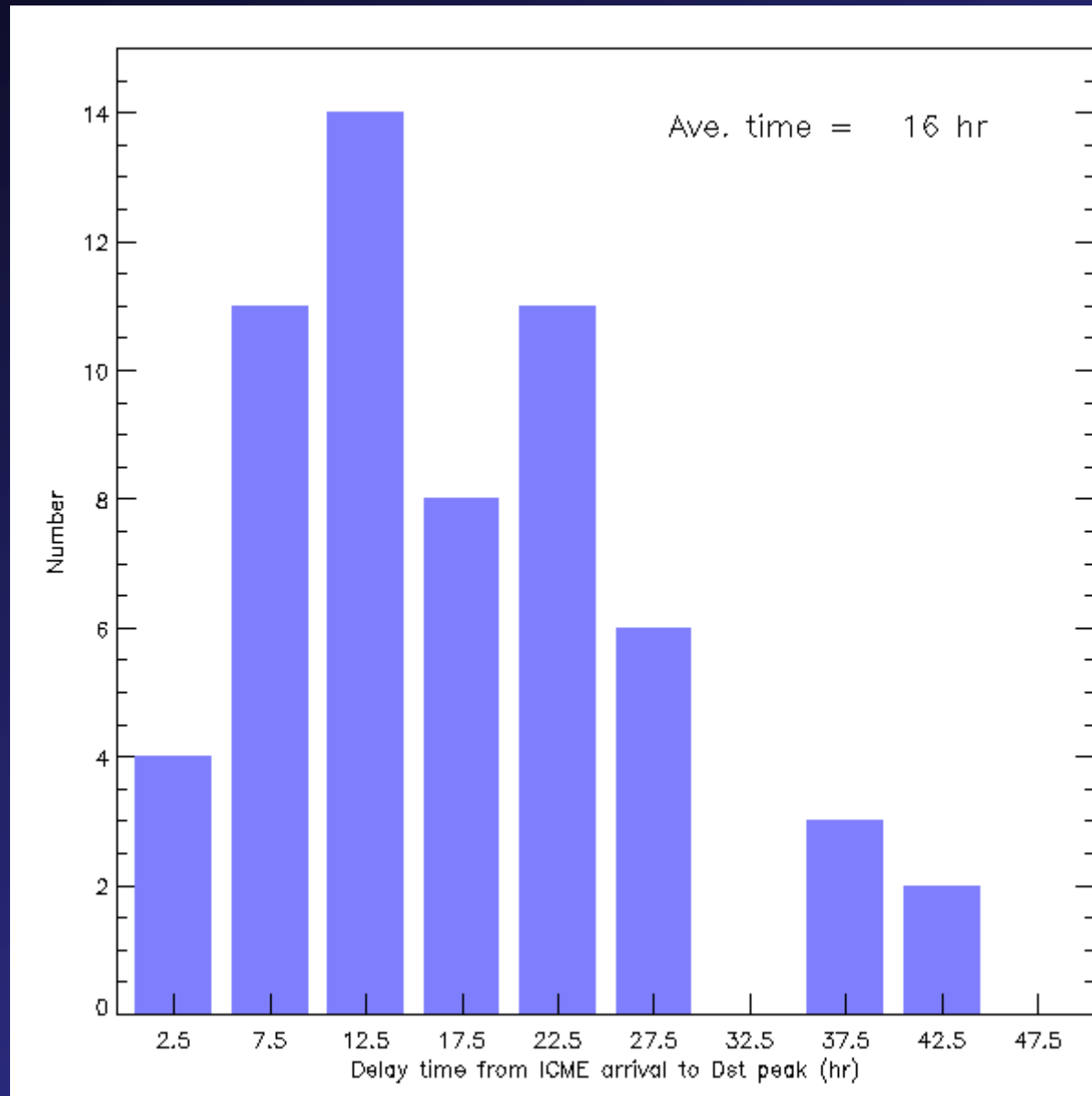
Velocity of CME and ICME & Transit Time



CME Transit Time from the Sun to Earth



Delay Time Betw. ICME Arrival and Dst Peak



Discussion and Conclusion

- **Three Types of Solar Drivers**

- S Type: 56%
- M Type: 26%
- C Type: 17%

- **Hemispheric Dependence on Solar Cycle**

- Western Hemisphere preferred for all phases
- Explanation: west hemisphere connection due to spiral IP field
- Northern Hemisphere preferred before polar field reversal
- Southern Hemisphere preferred after polar field reversal
- Explanation: Participation of global field in geo-effective Bz component
- North-Eastern quadrant is un-favored for all phases in solar cycle 23

- **Inner Heliospheric Observations are needed to identify the Sun-Earth-connection chain and predict the arrival time**